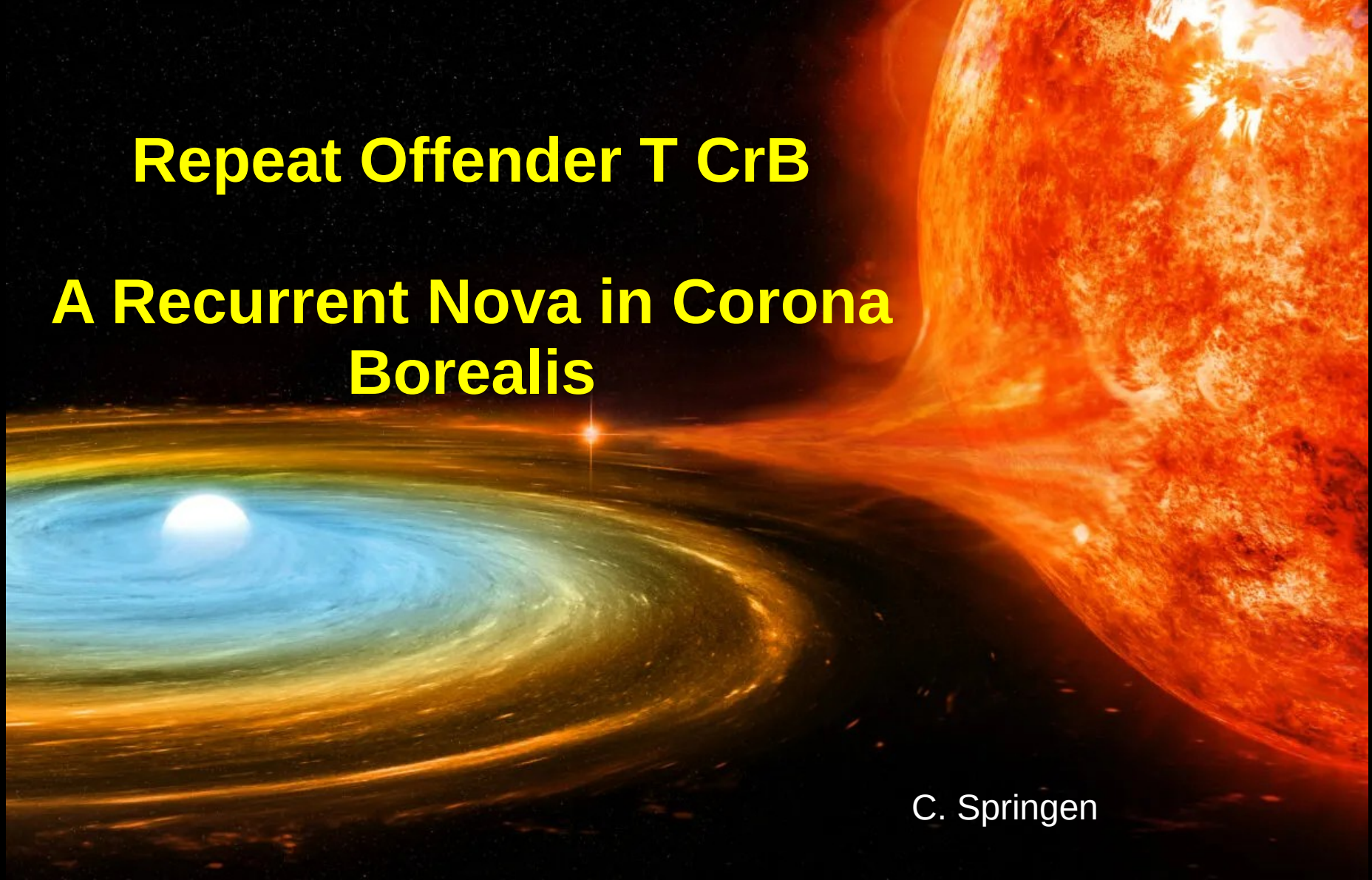


Repeat Offender T CrB

**A Recurrent Nova in Corona
Borealis**

C. Springen



T Corona Borealis

- Preliminaries – Terms & Jargon
- Progenitors – Red Giant & White Dwarf
- Binary Stars acting up - Stellar Bursts
- What am I supposed to do about it? Observing

Come to Terms with Astronomy

- Constellation names are in Latin
 - e.g. Coronae Borealis CrB (Northern Crown)
- Key symbols Earth = \oplus Sun = \odot
- pc parsec
 - 1 pc = 3.26 light-years = 206 k AU
- AU Astronomical Unit: mean Earth-Sun distance
 - 1 AU = 149 M km
- π parallax angle in arcseconds ["] or milliarcseconds [mas]
- M absolute & m observed magnitude (brightness)
- Luminosity = total power of EM radiation at all wavelengths (bolometric) emitted by a celestial object in ratio to the Sun L_{\odot} or in Watts

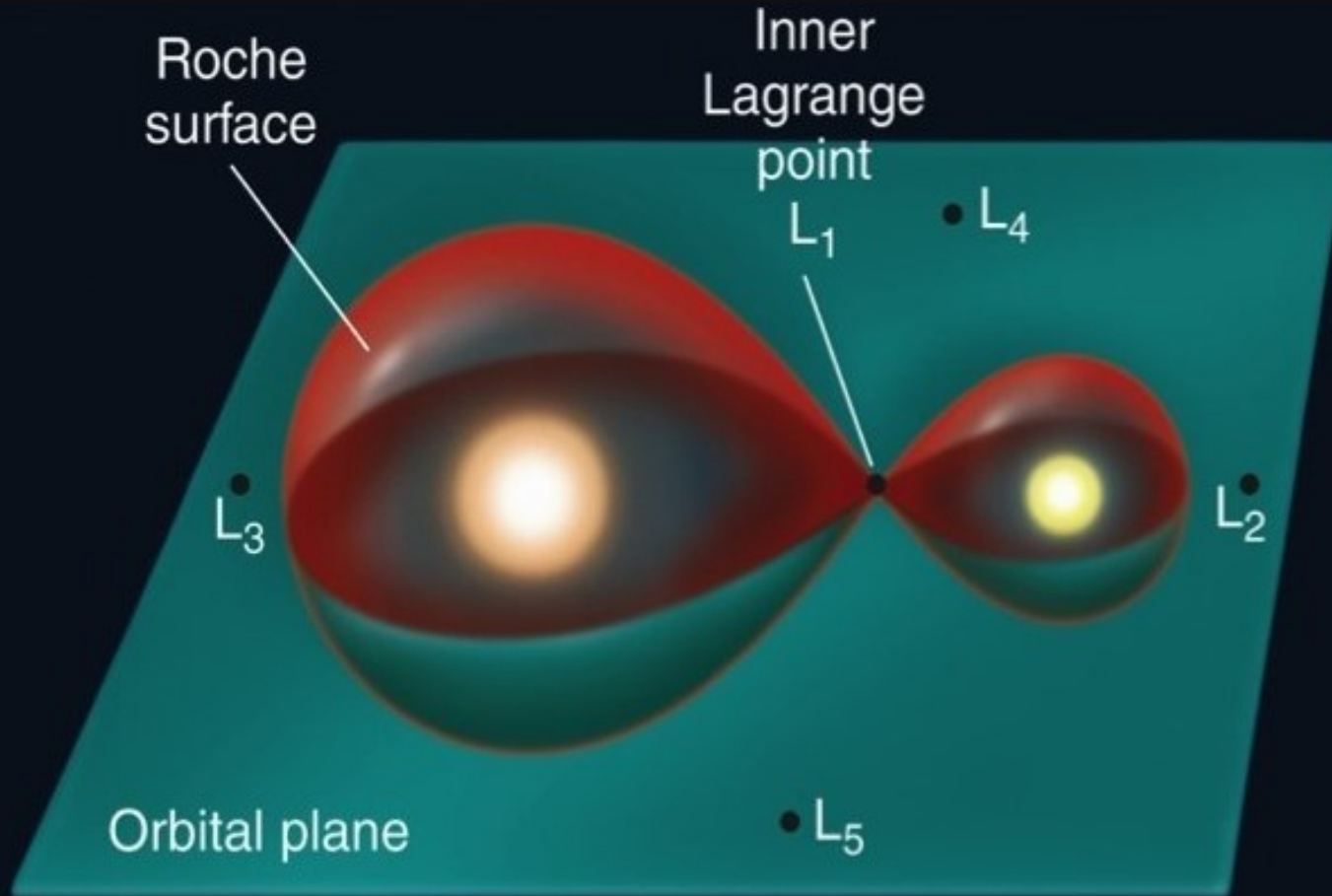
T CrB Description

- T CrB is a Binary system
 - Primary is a White Dwarf (WD)
 - Secondary (donor) is a M3III Red Giant
- T CrB is a Recurrent Nova (RNe)
 - Nova because of its dramatic brightening **& not** disintegrating its companion white dwarf (supernova).
 - Brightness expected to go from $m = 10$ binocular limit up to $m = 2.5$ naked eye visibility or 8 magnitudes
 - Recurrent with $p \approx 80$ years
 - Documented eruptions 1866, 1946 and Possibly 1217, 1787
 - Expected eruption 2025.5 +/- 1.3 yr (Schaefer 2023)
 - Precursor high-state occurred in 2015 (same as 1936)

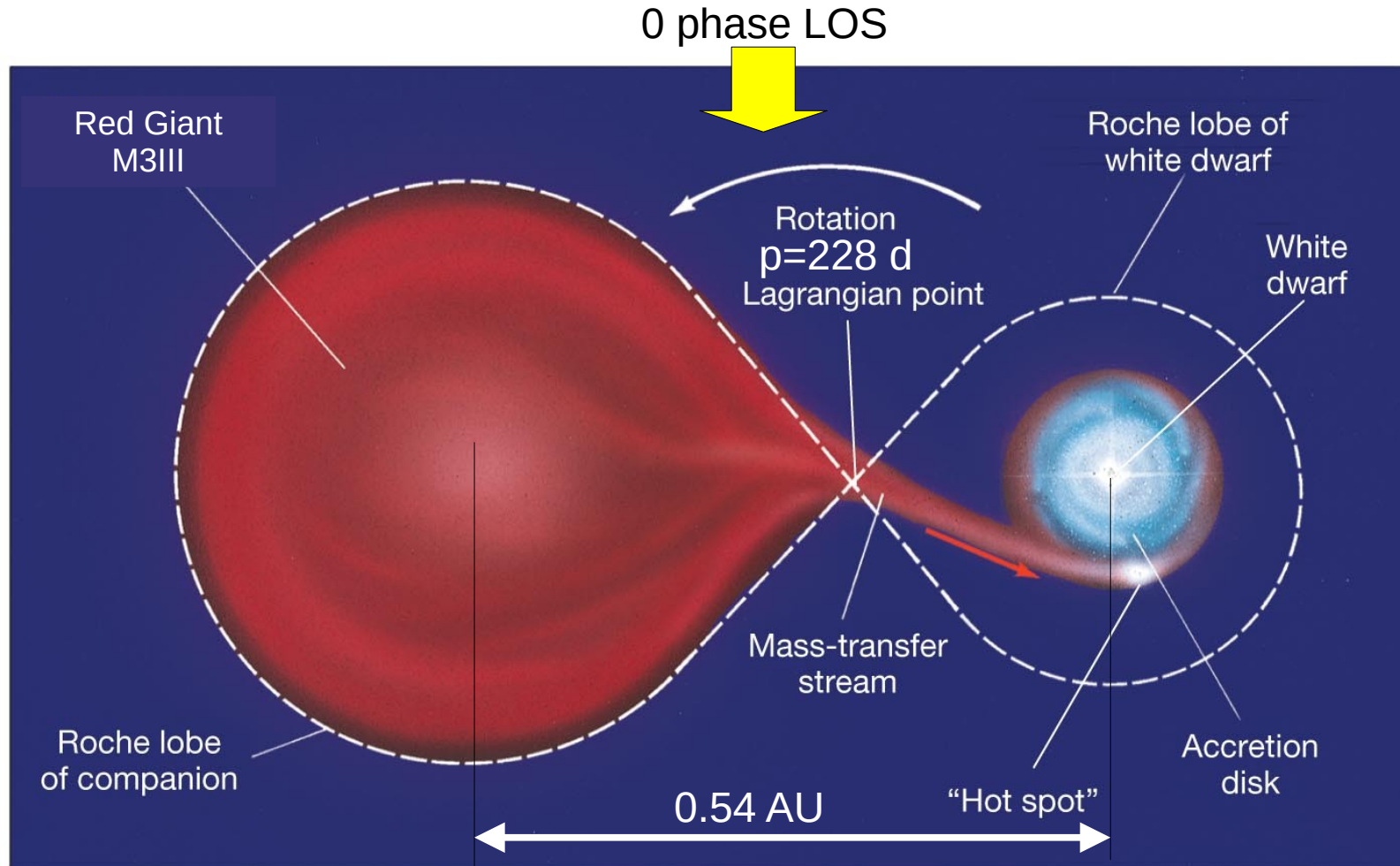
10 Recurrent Novae are Currently Known in the Milky Way Galaxy

- | | | | | | |
|---|-----------|------------|----|-----------------------------------|-------------------------------------|
| 1 | T Pyx | Pryxis | 6 | V394 CrA | Corona Australis |
| 2 | IM Nor | Norma | 7 | T CrB | Brightest of RNe |
| 3 | CI Aql | Aquila | | $m_{qV} = 10.1$ | $m_{peakV} = 2.5$ |
| 4 | V2487 Oph | Orphiuchus | 8 | RS Oph | (Erupted Aug 2021) |
| 5 | U Sco | Scorpius | 9 | V745 Sco | Scorpius |
| | | | 10 | V3890 Sgr | Sagittarius |

Teardrops of a Stellar Binary

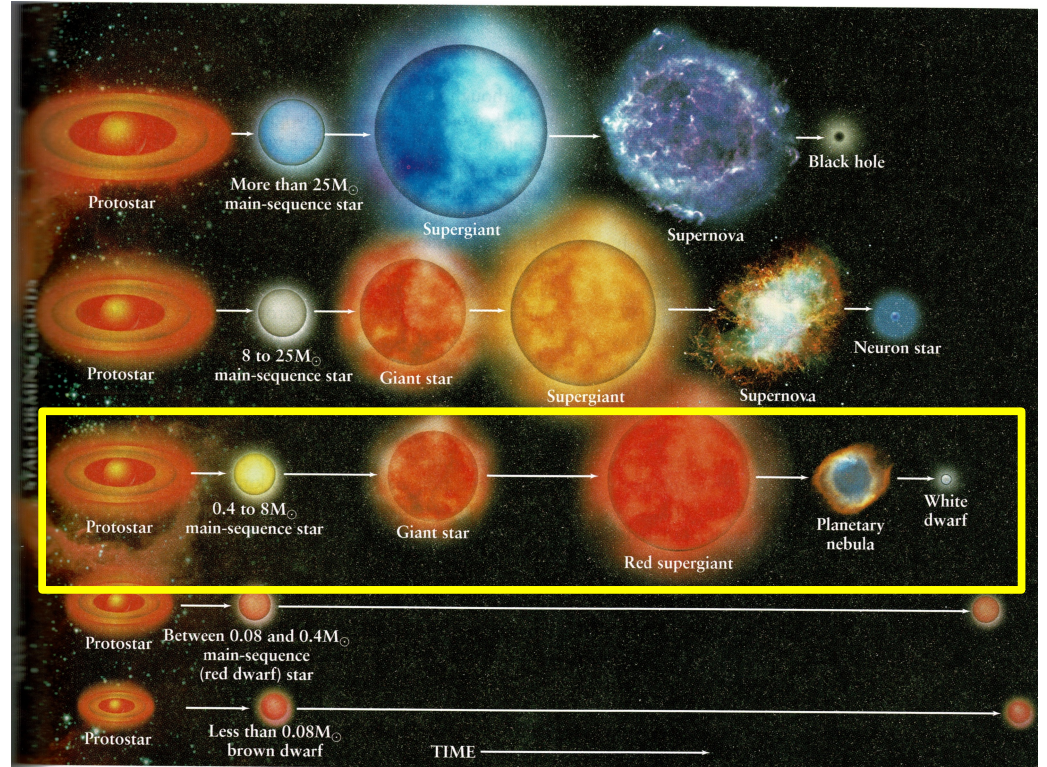


T CrB Nova Binary System



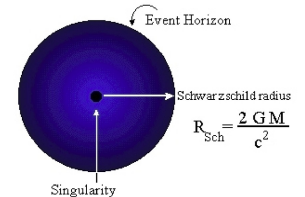
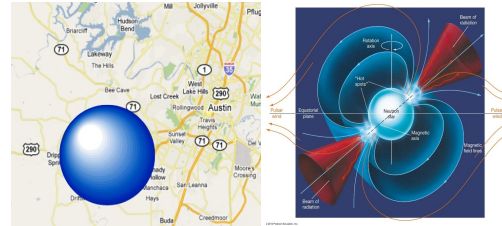
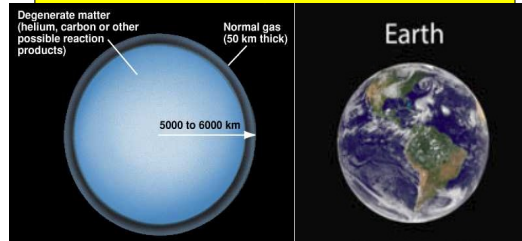
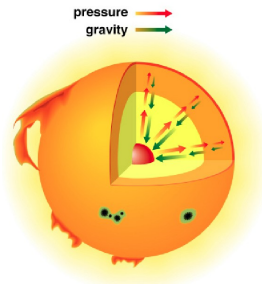
Evolution of Stellar Objects

- T CrB - Primary - White dwarf
 - Progenitor is main sequence Sun-like star $0.4 - 8 M_{\odot}$
 - End state is electron degenerate WD $M = 1.37 M_{\odot}$
 - Dense M_{\odot}/V_{\oplus}
- T CrB - Secondary Red Giant
 - Evolved Sun-like star
 - $M_{\text{RG-TCrB}} = 1.12 M_{\odot}$



Equilibrium - the Balance of Forces

Ideal Gas Law	Pauli Exclusion Principle		TBD Physics
Hydrostatic Equilibrium	Electron Degeneracy	Neutron Degeneracy	Singularity Quark degeneracy?
Main sequence & giant stars	White Dwarfs	Neutron Stars	Black Holes
$R = .1 - 1500 R_{\odot}$ $M \approx 0.1 - 200 M_{\odot}$ $\rho \approx 10^3 \text{ kg/m}^3$	$R \approx 6 \times 10^3 \text{ km}$ $M \approx 1 M_{\odot}$ $\rho = 10^9 \text{ kg/m}^3$	$R \approx 10 \text{ km}$ $1.4 < M < 3 M_{\odot}$ $\rho = 10^{17} \text{ kg/m}^3$	$R_s = 3 * M \text{ km}$ $3 < M < 10^9 M_{\odot}$ $\rho = \infty$



T CrB System Physical Data

- Binary Star M3III Red Giant + WD
 - M3 Red Giant, cool (3600 K)
 - $M_{\text{TM3}} = 1.12 M_{\odot}$
 - $R_{\text{TM3III}} = 75 R_{\odot}$
 - White Dwarf, hot (10,000 K)?
 - $M_{\text{TWd}} = 1.37 M_{\odot}$ Chandrasekhar limit = $1.4 M_{\odot}$
 - $L_{\text{TWd}} = 100 L_{\odot}$
- Separation is $d = .54 \text{ AU}$
- Orbital Period = 228 d

Implications of Physical Parameters

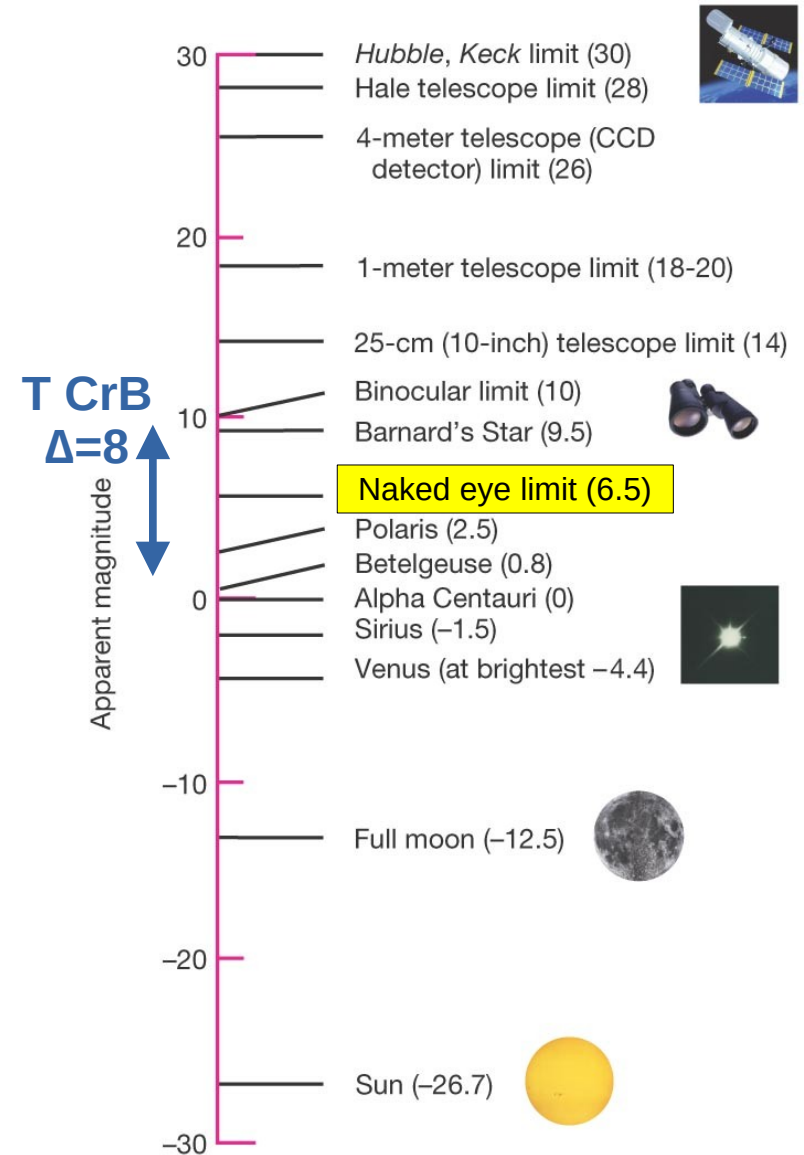
- Recurrence rates vary from 10 - ~100,00 yr
- Fast recurrent RN_e i.e. $\tau_{\text{rec}} < 100$ yr imply
 - Mass of primary is close to Chandrasekhar mass $M_{\text{CS}} = 1.4 M_{\odot}$ $M_{\text{TCrB-WD}} > 1.37 M_{\odot} \checkmark$
 - Sufficient mass transfer rate to primary $\sim 10^{-7} M_{\odot} \text{yr}^{-1}$
- Mass of evolved secondary Red Giant over fills the Roche Lobe and spills onto the primary White Dwarf
- When sufficient Hydrogen builds up thermo-nuclear fusion initiates and causes major brightening of system

Some Research Results

- Hydrogen is accreting onto surface of the WD
- Base layer of H is electron degenerate so cannot expand thermally
- T CrB WD thermonuclear runaway cycle is imminent (2024-2025) resulting in outburst
- Evidence suggest a net accretion of material
 - Secular evolution of the WD may allow it to reach to the Chandrasekhar limit and explode as either a Type Ia supernova or undergo accretion induced collapse, depending on the composition of the WD.

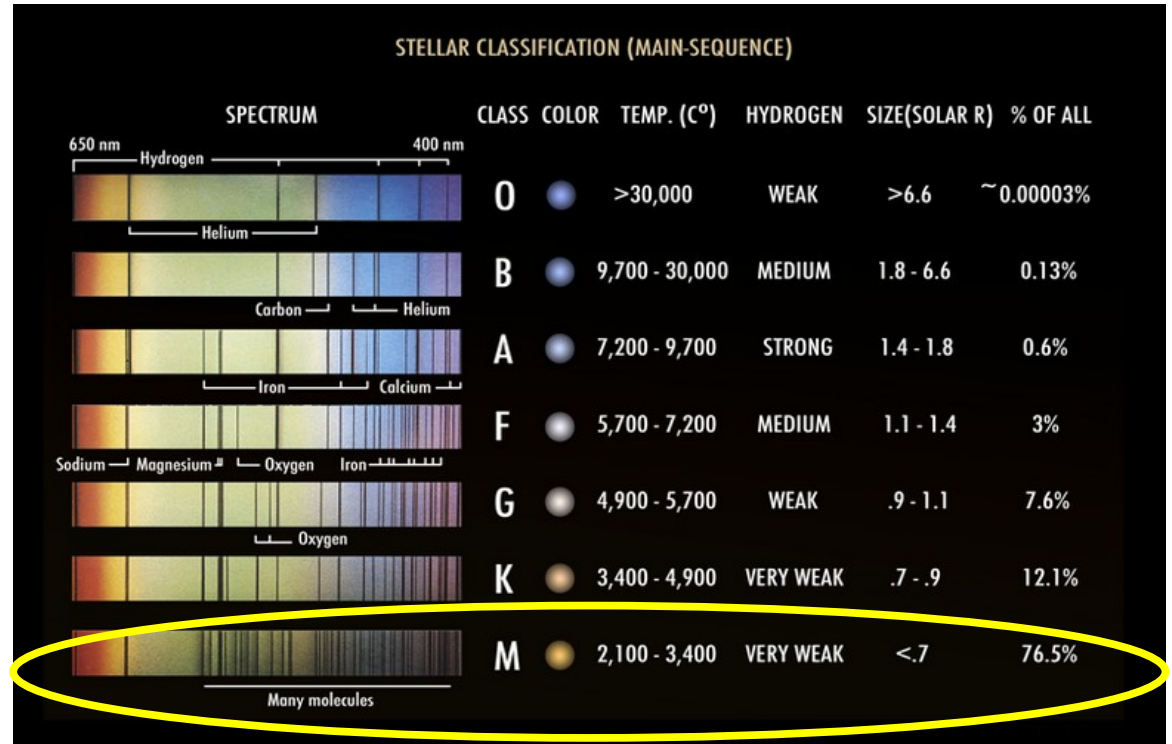
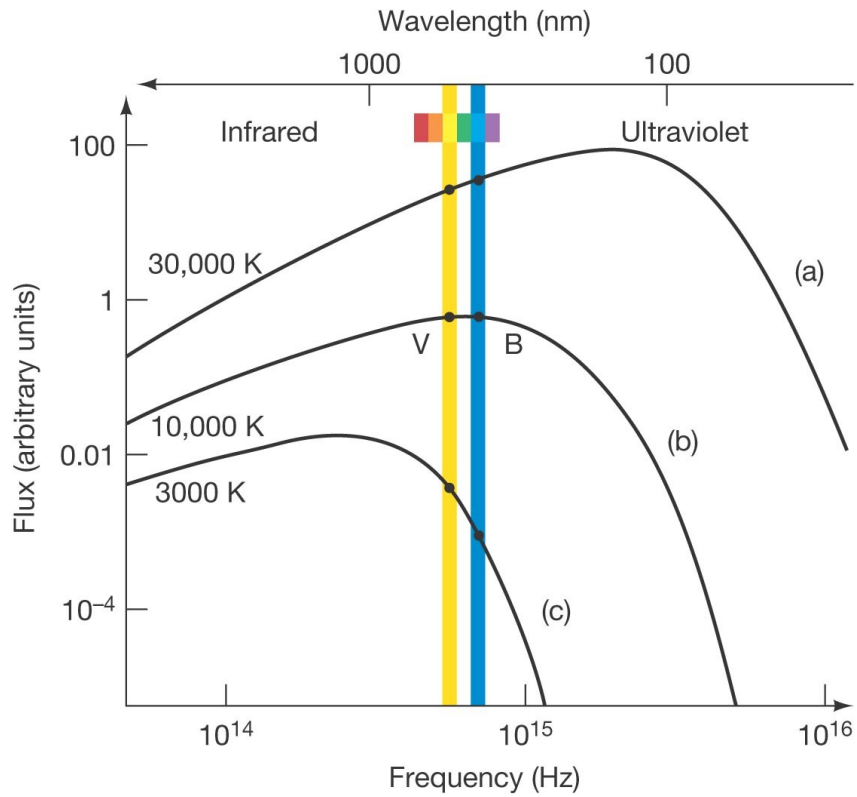
Brightness

- Measured in Magnitudes
 - Absolute magnitude **M** is @ 10 pc
 - Apparent magnitude **m** is as observed locally
 - Inverted scale – more positive is dimmer
 - Logarithmic scale
 - 5 Magnitudes is 100 x in brightness
 - $m \approx -2.512 \log_{10} (F_x/F_{ref})$ F_{ref} w.r.t Vega
 - Example: Sun $m = -26.7$ & $M = 4.8$



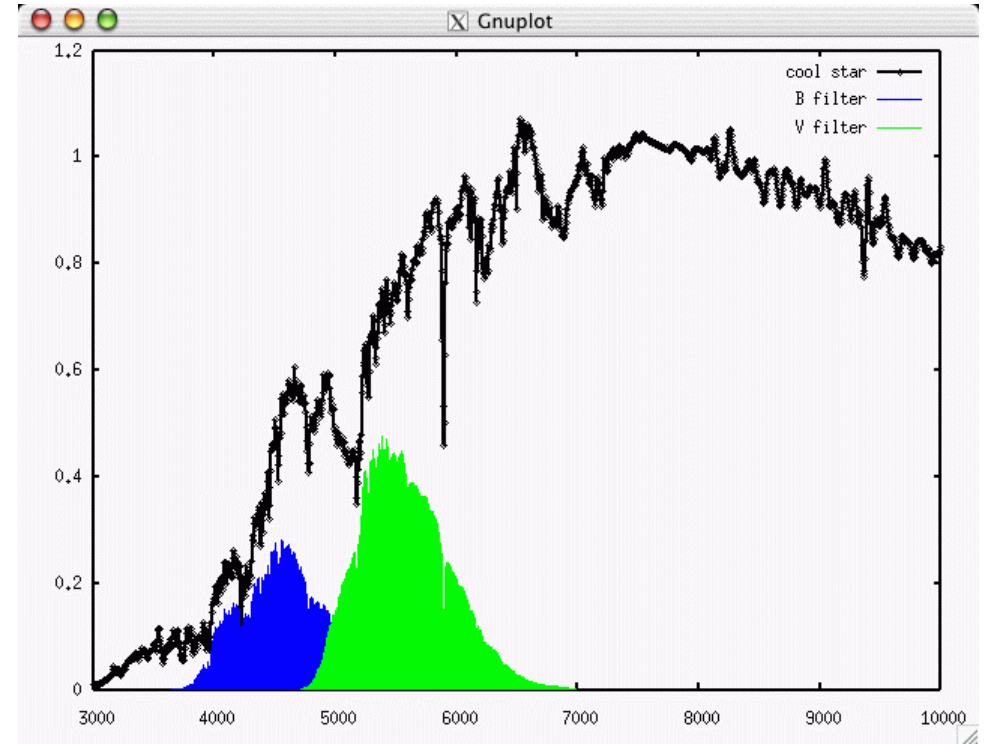
M Spectral Type

- Cool star $M < 3400$ K



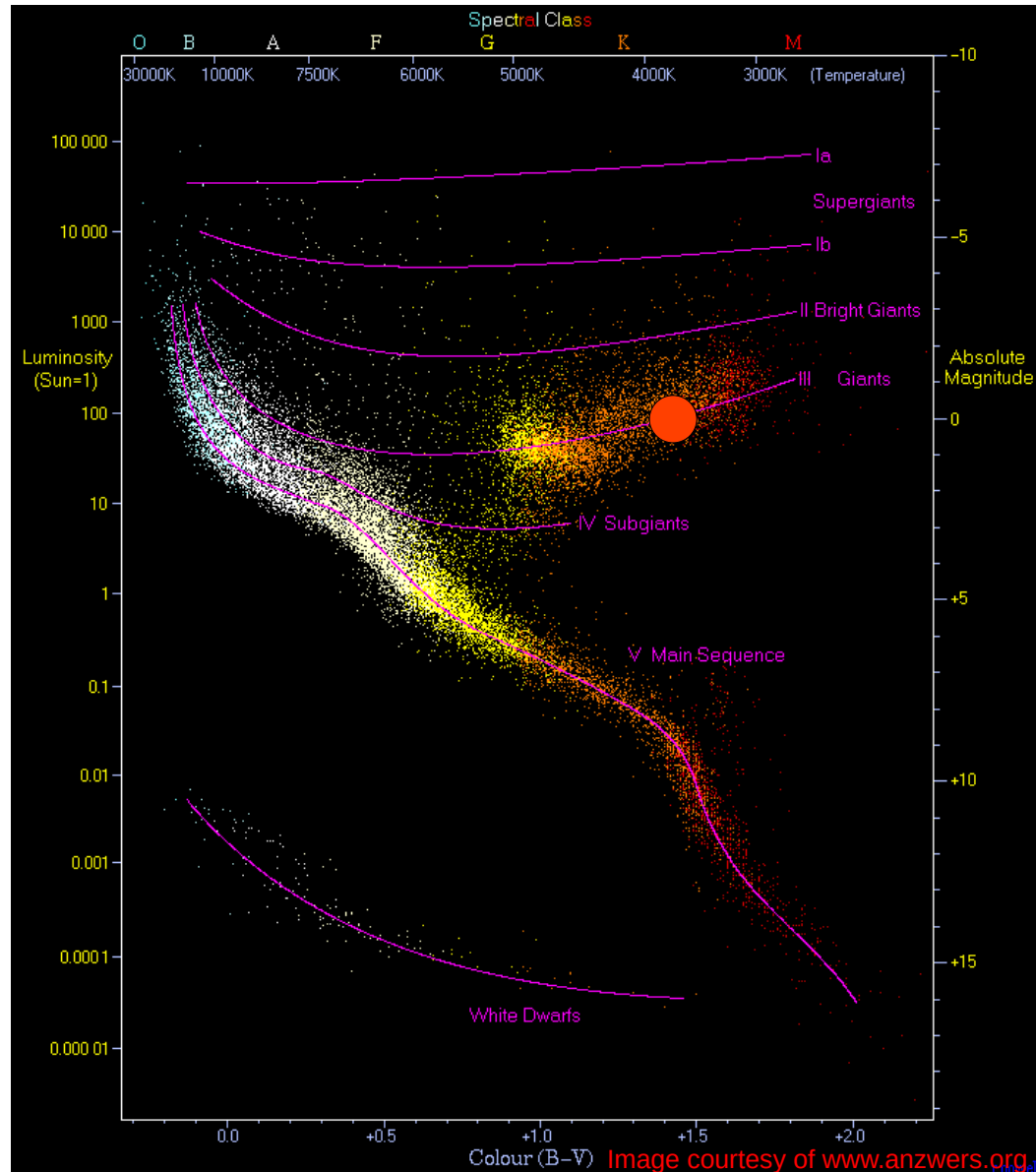
Temperature via Color Index B-V

- T CrB Quiescent color ratio
 - $B_q - V_q = 11.6 - 10.1 = 1.5$
(very red - cool)
- Peak color ratio (in outburst)
 - $B_{pk} - V_{pk} = 2.5 - 2.5 = 0$
(hotter - like Vega)



Luminosity Class

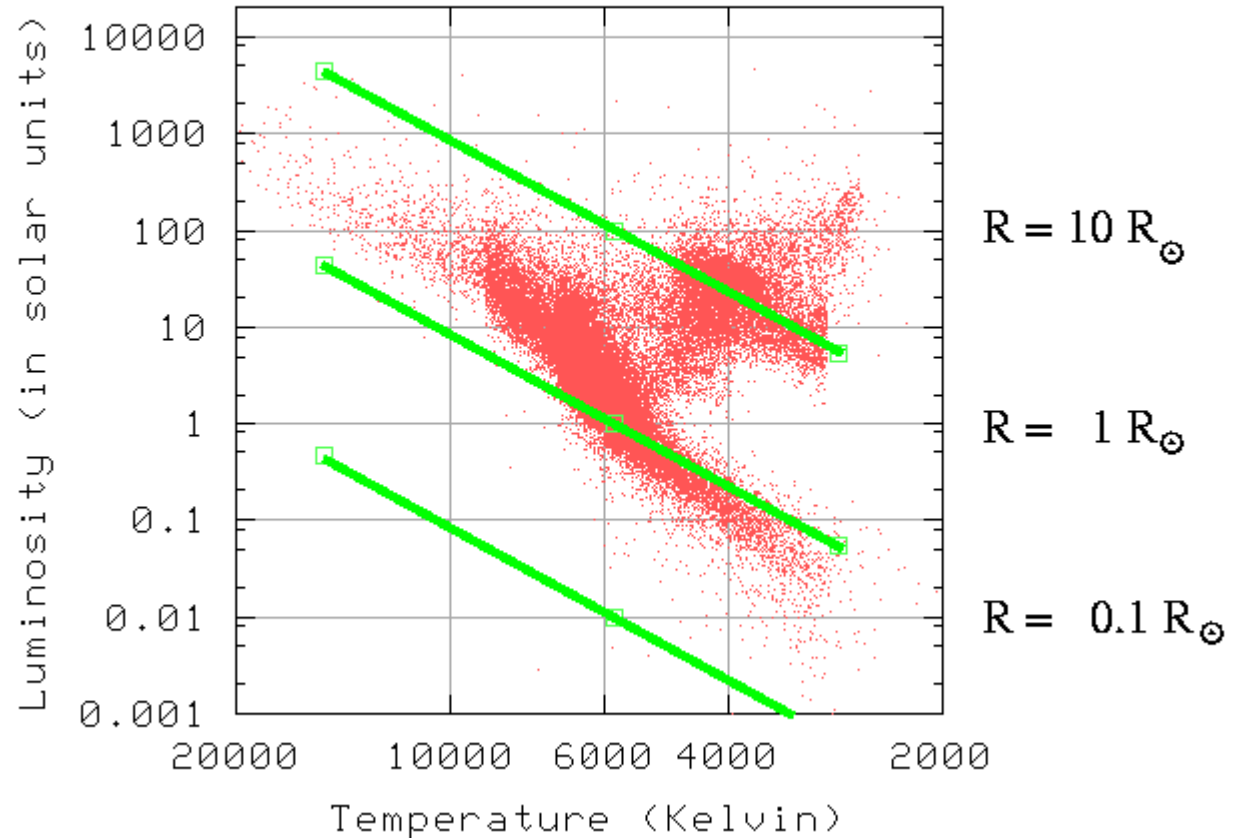
- Determined from spectral line widths
- T CrB Primary is a white dwarf
 - Stellar core
 - No fusion → cooling
 - Very small → low luminosity
- T CrB Secondary is a luminosity class III evolved *Red Giant* star
- You only get dwarfs and Giants no *normal* stars :-)



Size on a HR Diagram

- Diagonal Lines plot constant radii
- Main Sequence stars are about $1 R_{\odot}$
 - OB stars a little more
 - M stars a little less
- Note spacing is logarithmic

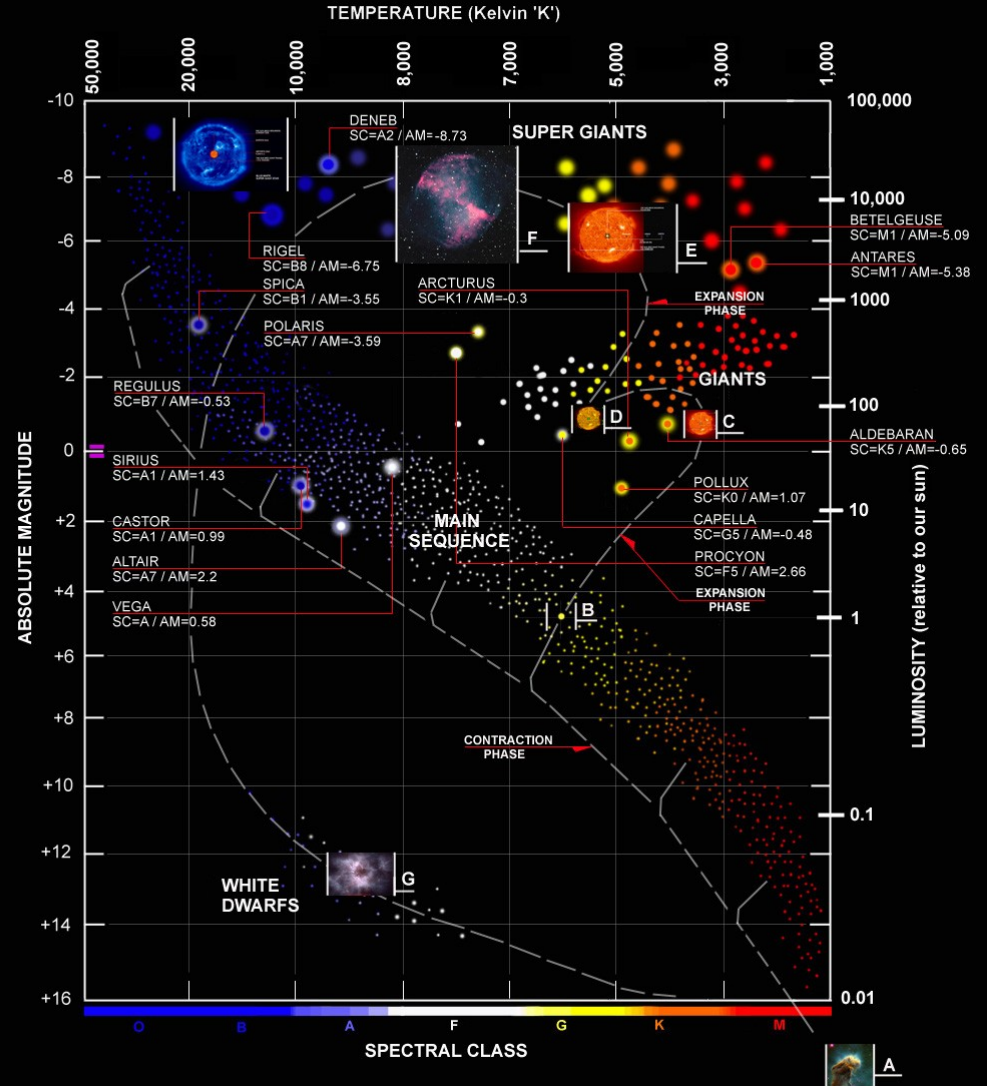
HR diagram of nearby stars



Hertzprung-Russel (HR) Diagram

Classification

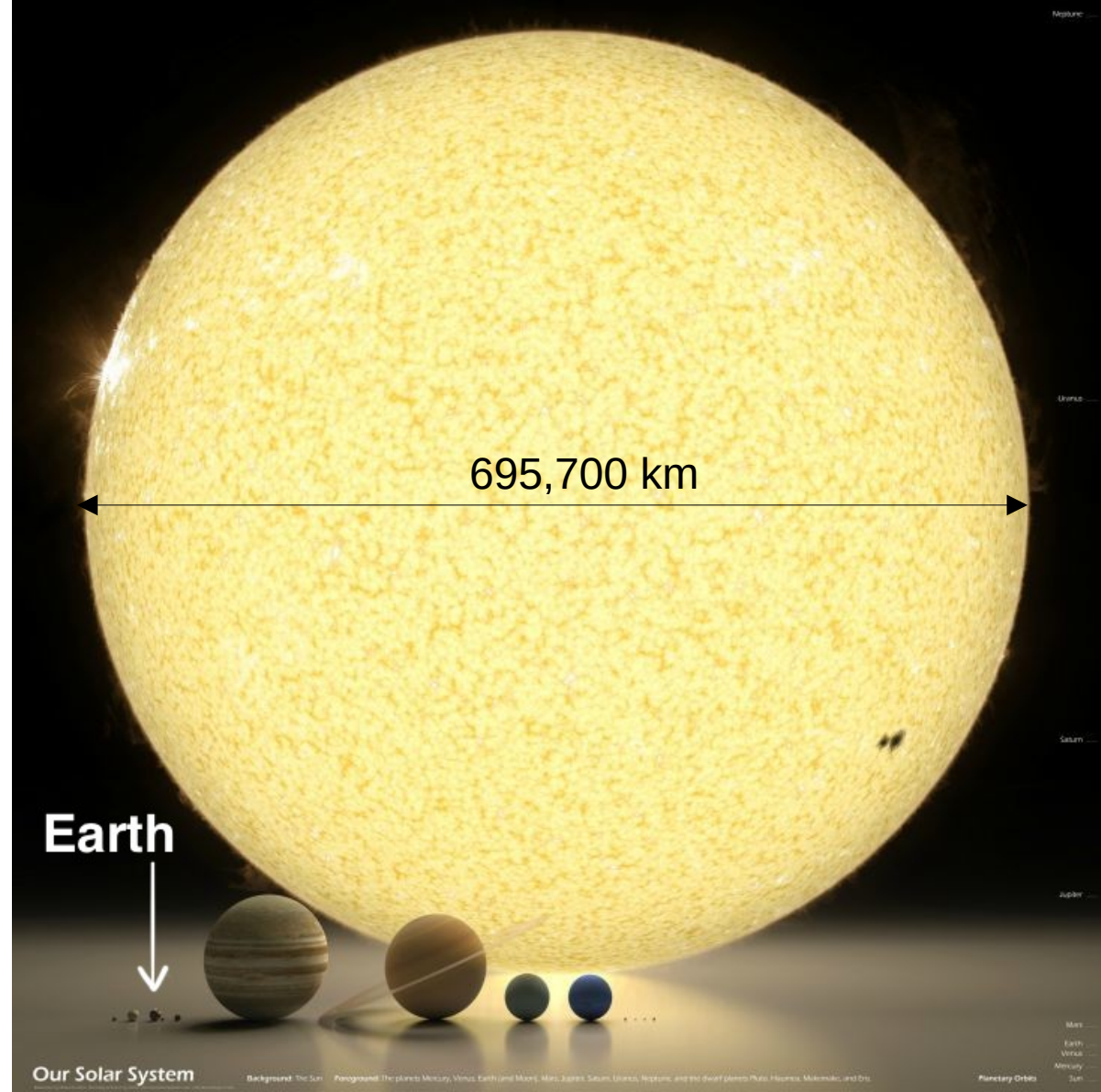
- Evolution
- Size
- Temperature
- Brightness/Luminosity
- Main Seq. Mass



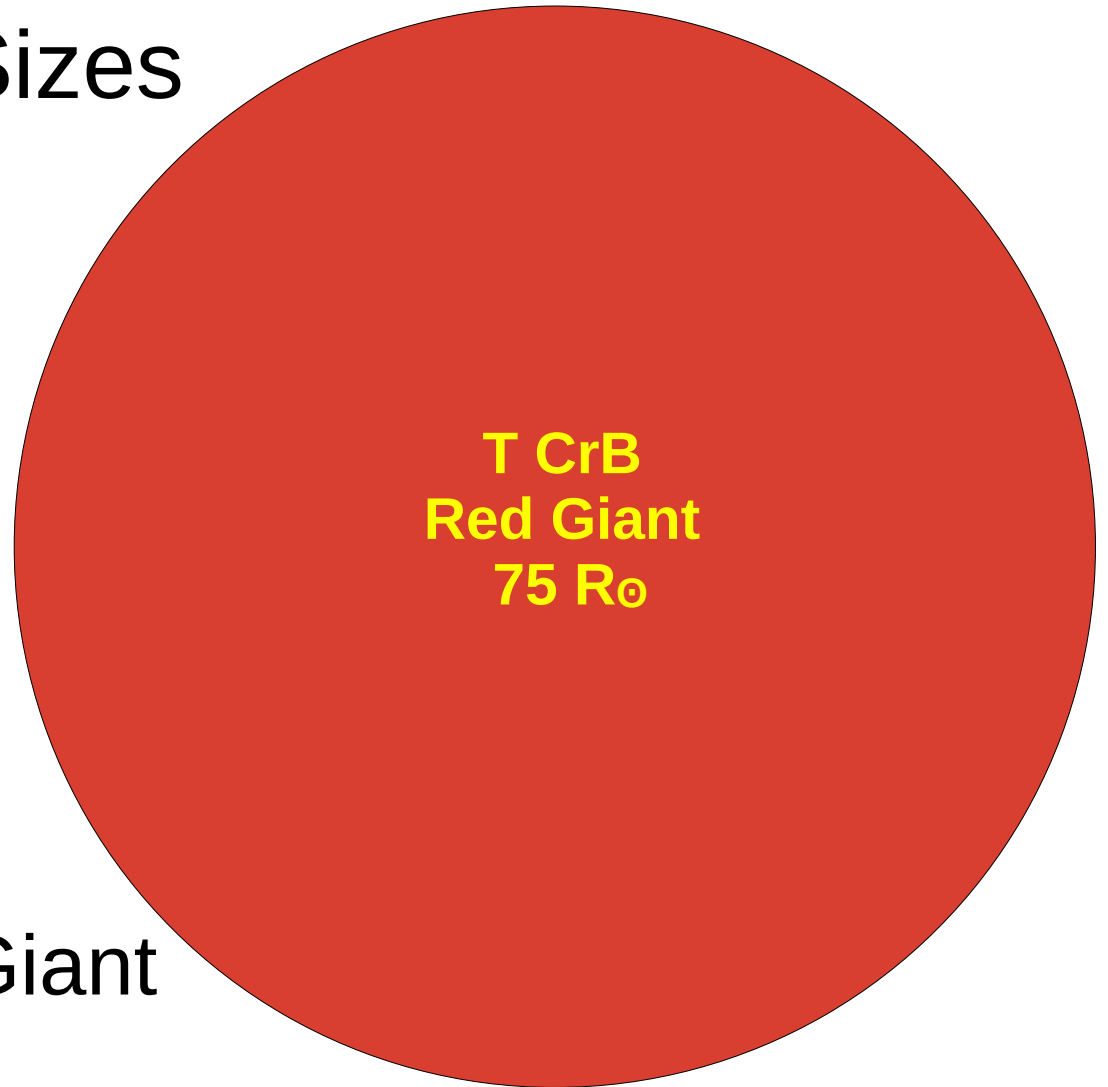
Our Sun's Size relative to Earth

$$R_{\odot} = 109 R_{\oplus}$$

T CrB White Dwarf is about the size of Earth.



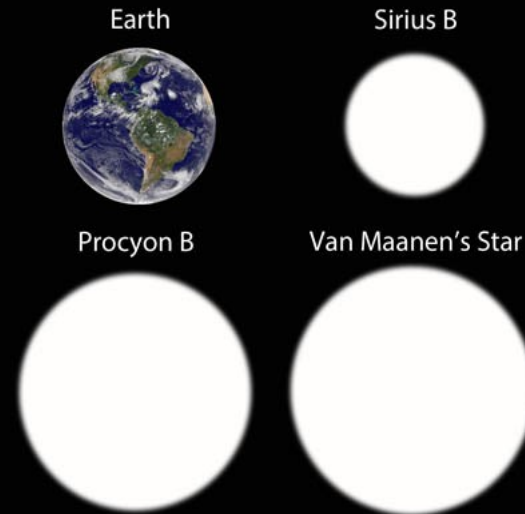
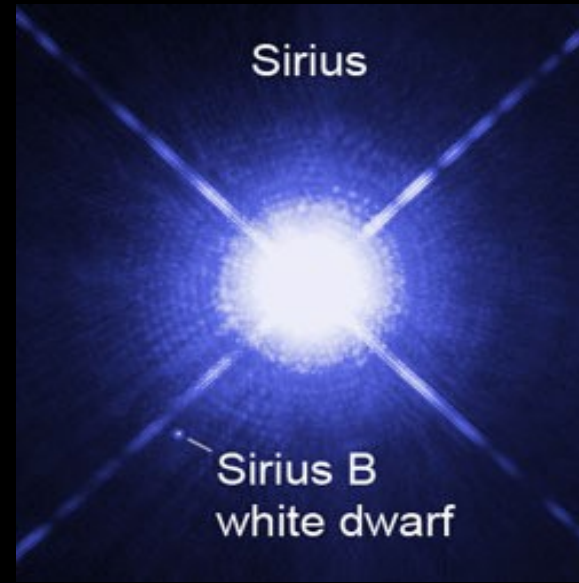
Sun & M3III Giant Sizes Relative to Sun



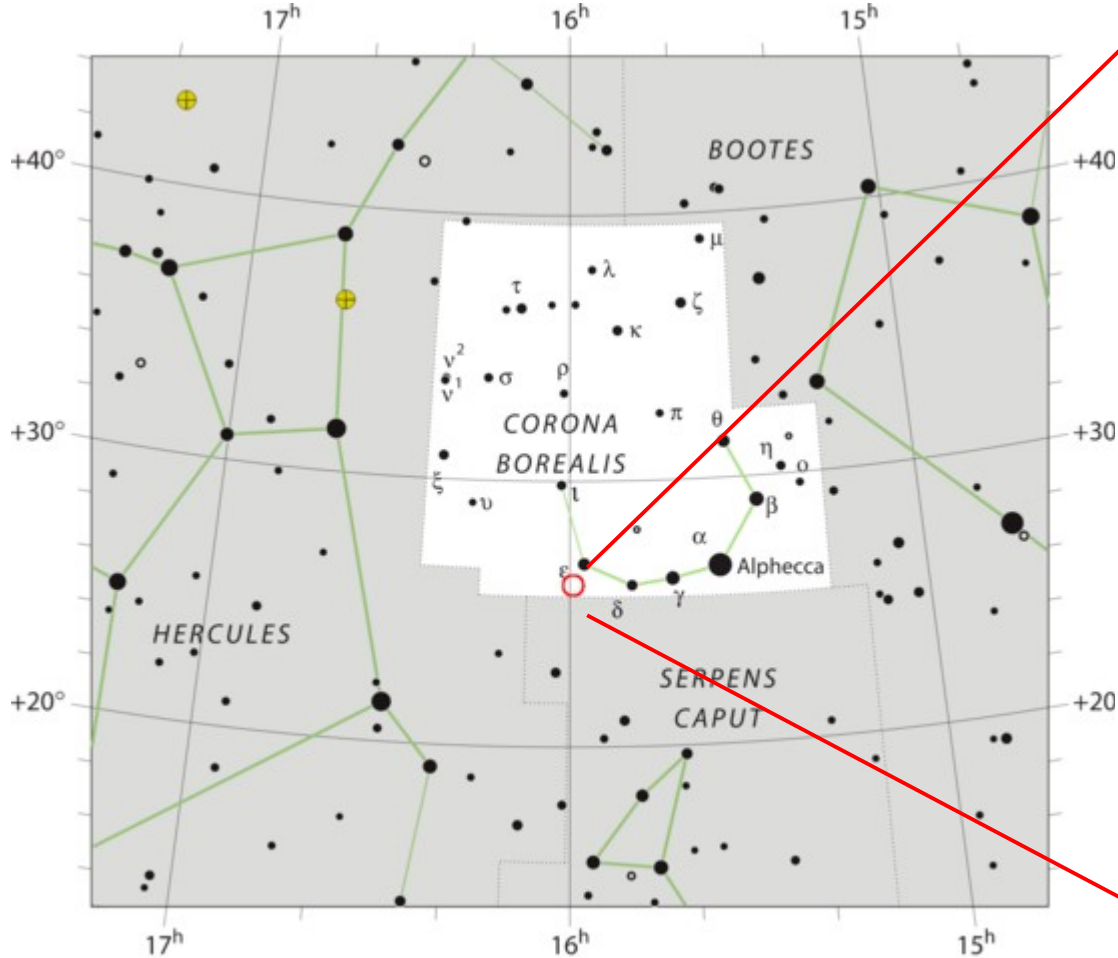
- Companion is a luminosity class III Giant

White Dwarfs

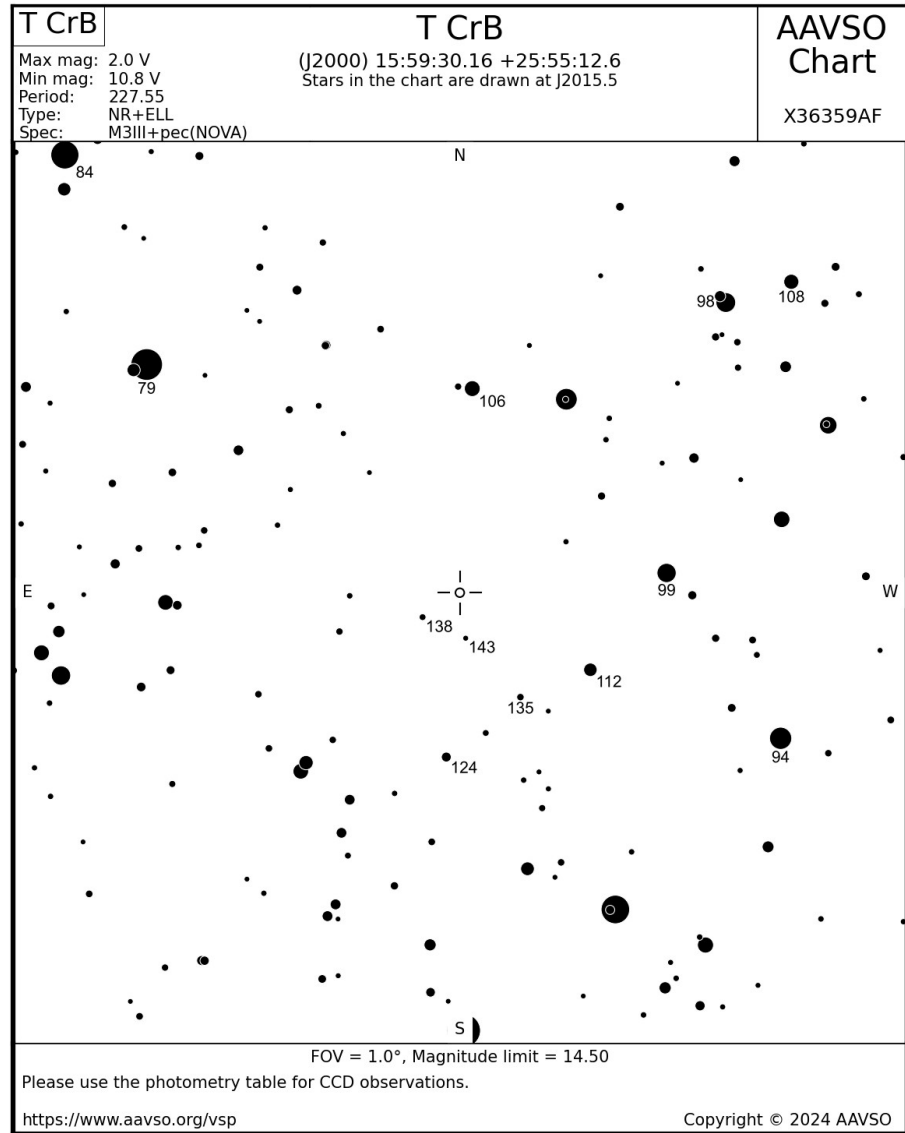
- Small, Earth sized +/-
- Massive $\sim M_{\odot}$
- Dense $\sim 10^9 \text{ kg/m}^3$
- Electron degenerate
- No fusion
- Initially hot (white) then cool to Black Dwarf & crystallizes



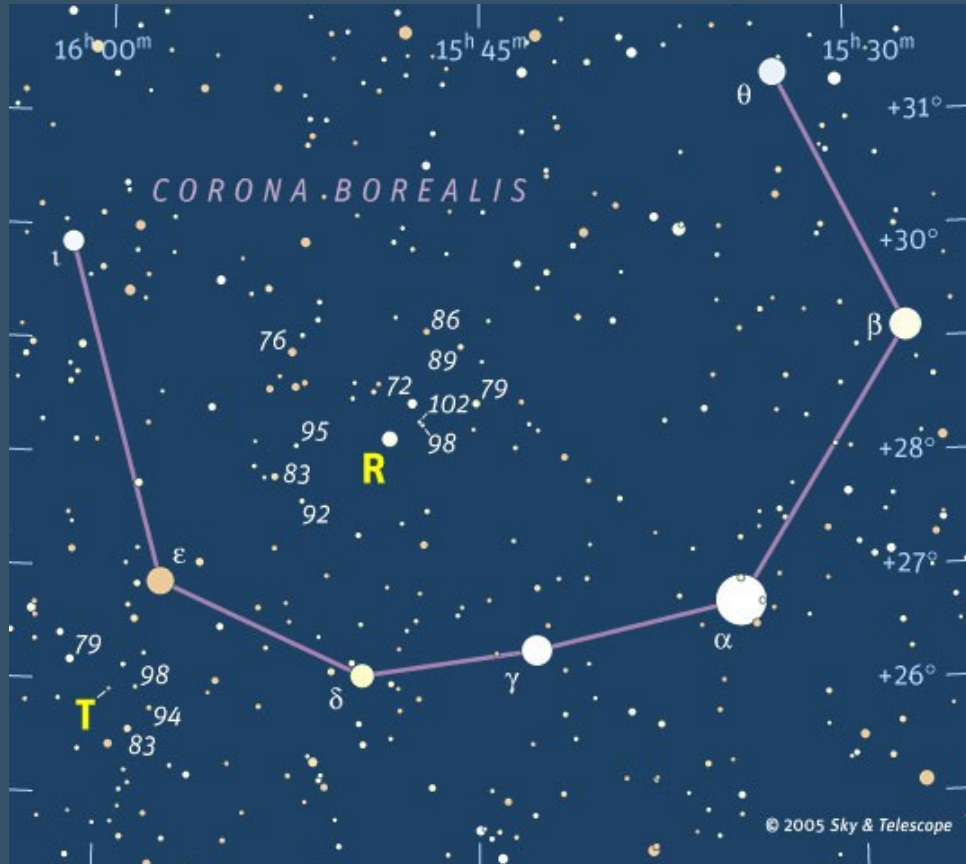
Finder Charts



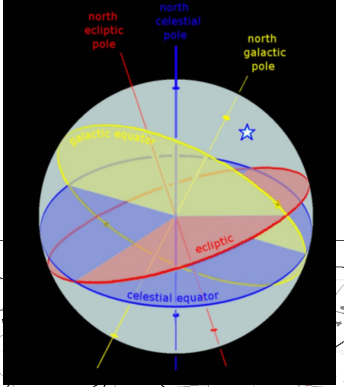
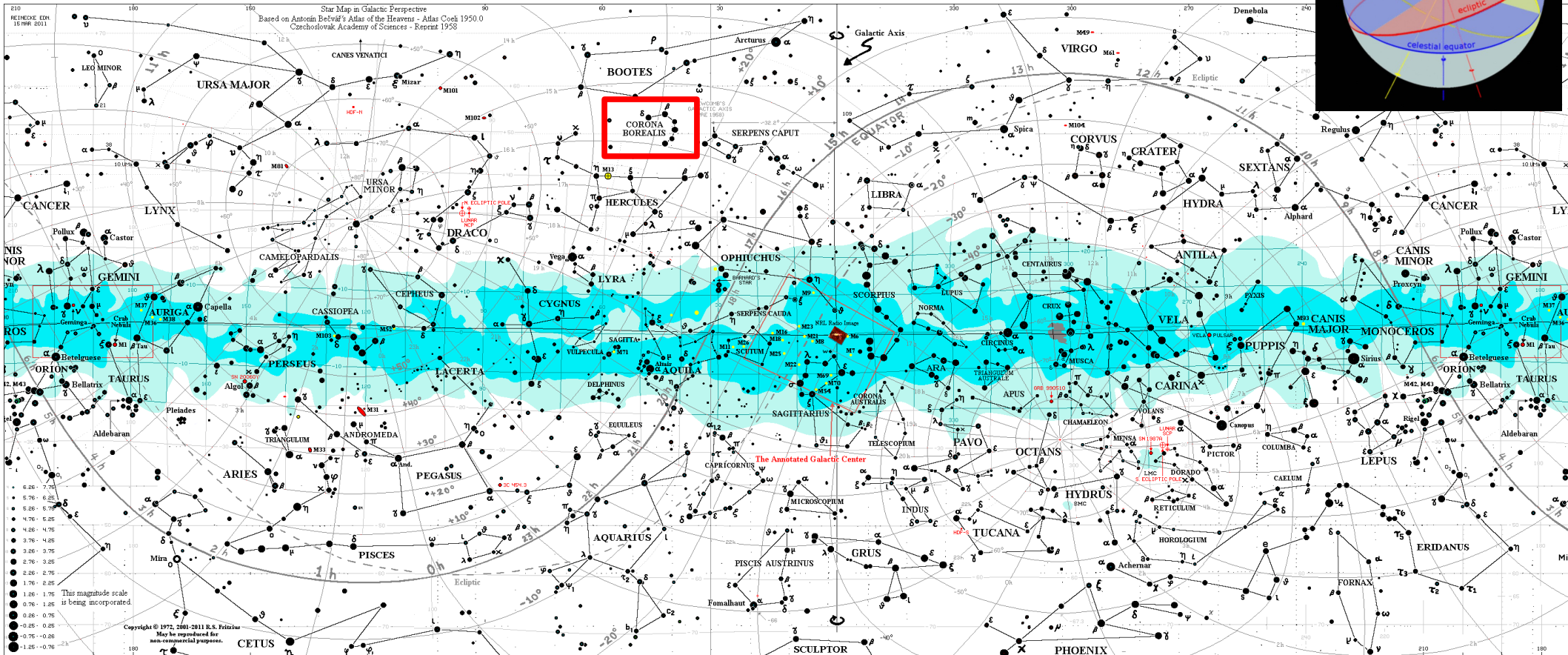
● 2 ● 3 ● 4 ● 5 ● 6



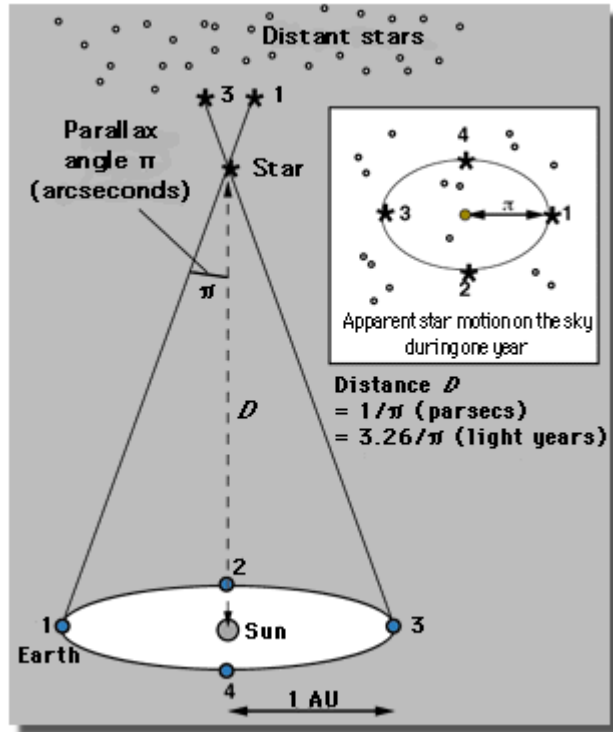
T CrB is Less Than Prominent in Quiescence



Galactic Projection with T CrB



Parallax Distance of T CrB



- T CrB parallax via Gaia Space Telescope
 - $\pi = 1.1538$ mas
 - $d = 1/p = 1/(1.1538 \times 10^{-3} \text{ as}) = 866.7$ pc
 - ± 17.7 pc
 - $d \approx 867$ pc

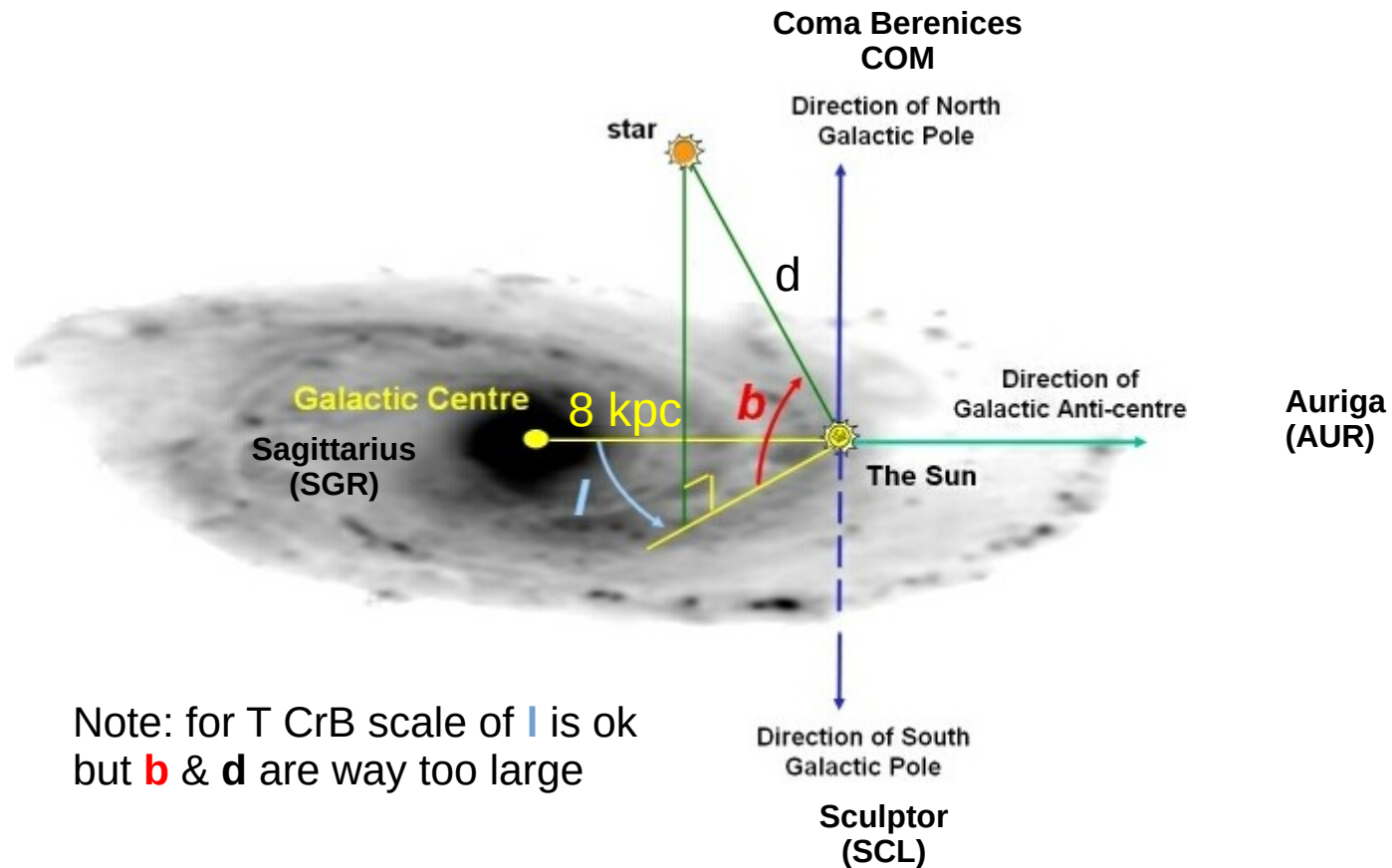
Gaia DR3 4111834567779557376

Astrometry	Photometry	Spectroscopy	Astrophysical parameters
Description	Value		Unit
Equatorial ICRS (RA,DEC) at epoch 2016	256.5229102004, -26.5805651308		deg
Galactic (l, b) at epoch 2016	357.0803450631, 8.5731964881		deg
Parallax	1.1538 ± 0.0241		mas
RA proper motion	0.3896 ± 0.0256		mas yr ⁻¹
DEC proper motion	-0.2893 ± 0.0165		mas yr ⁻¹
Renormalised unit weight error	0.837		

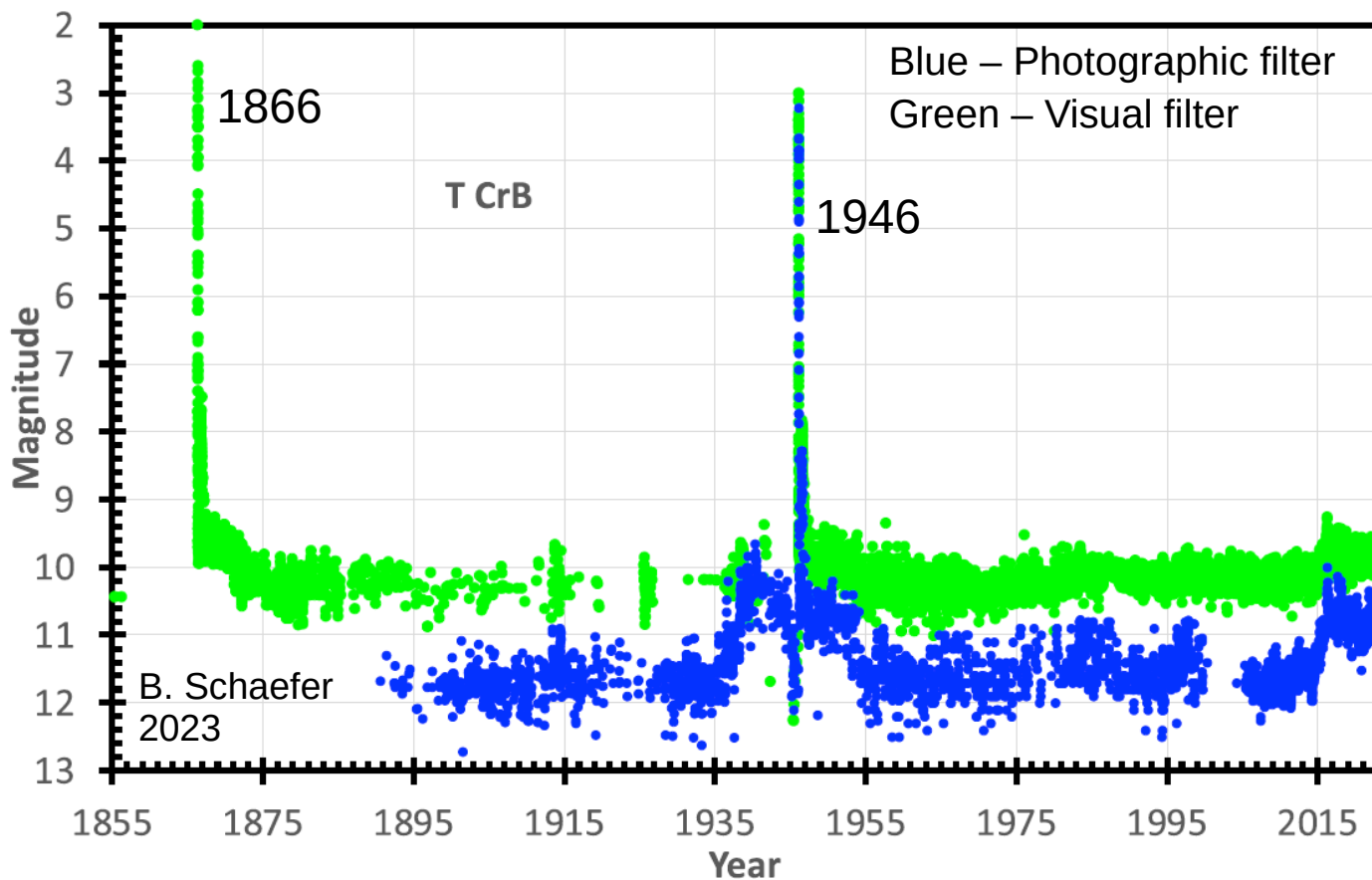
Galactic Coordinates

T CrB

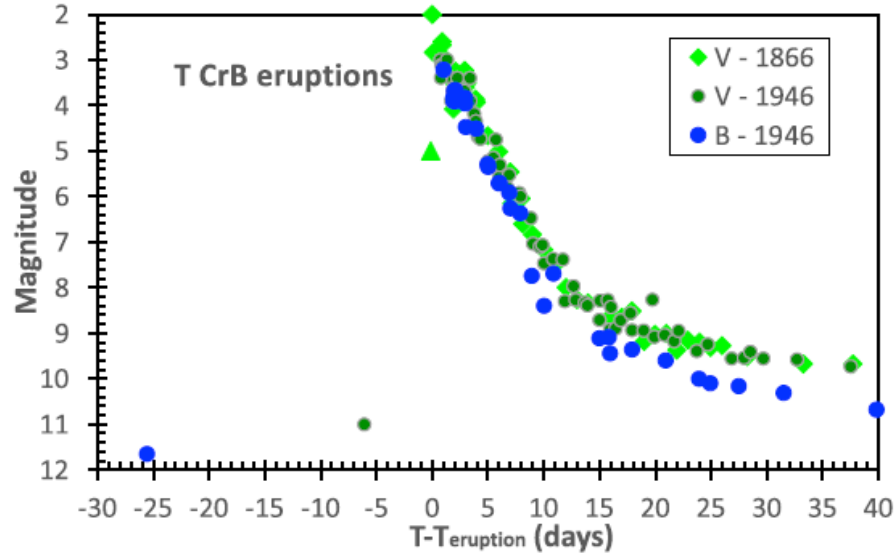
- $b = 48.1648^\circ$
- $l = 42.3738^\circ$
- $d = 0.9 \text{ kpc}$



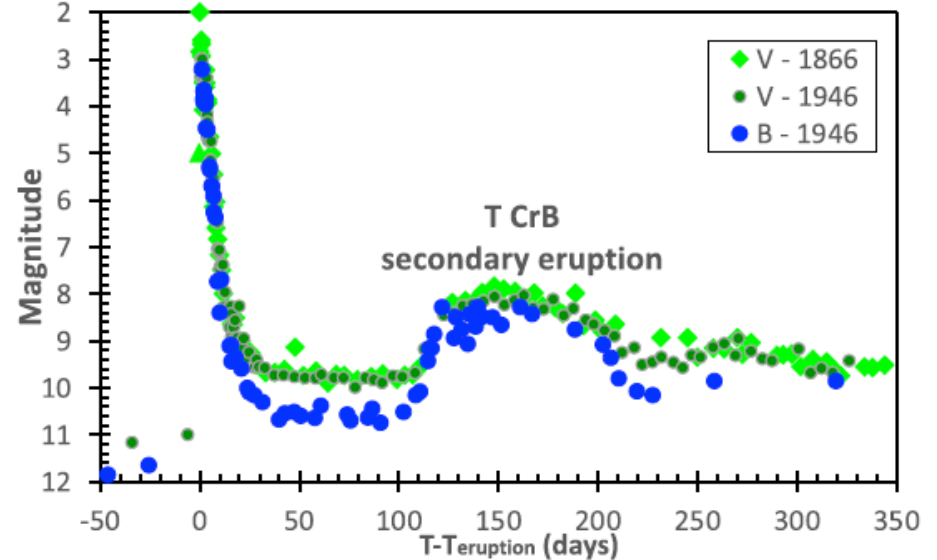
Observed Outbursts of T CrB



B & V Band Light Curves

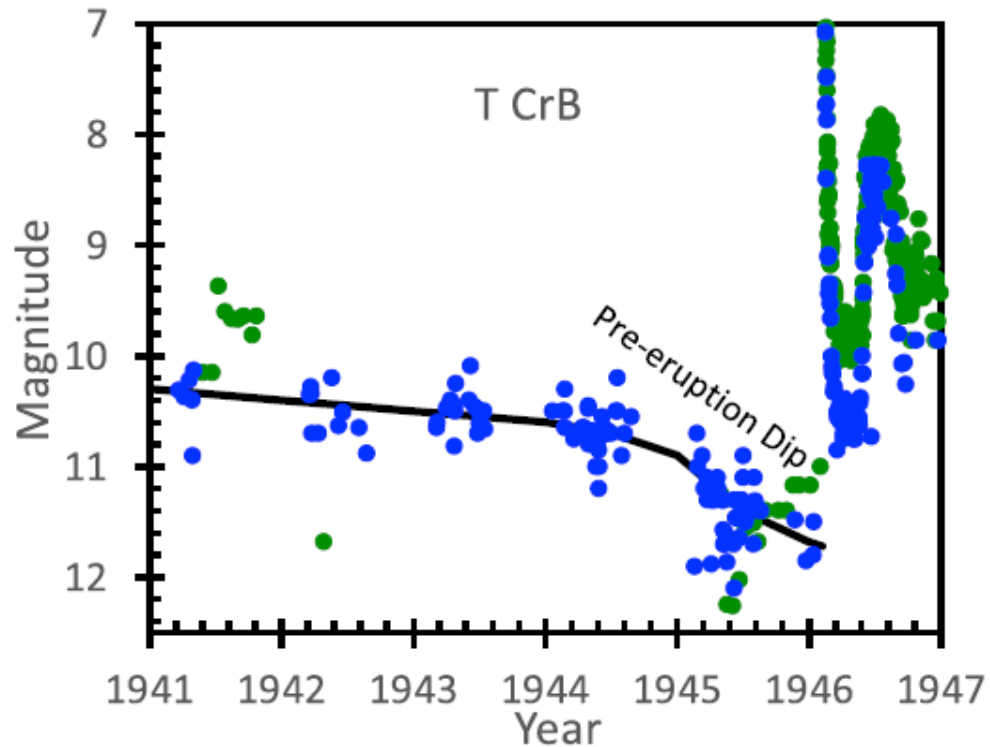


- Rise time $\sim 1 \text{ Mag hr}^{-1}$
- Fall 3 Mag 5 days



- Unusual for a Nova - secondary eruption
 - Explanation TBD

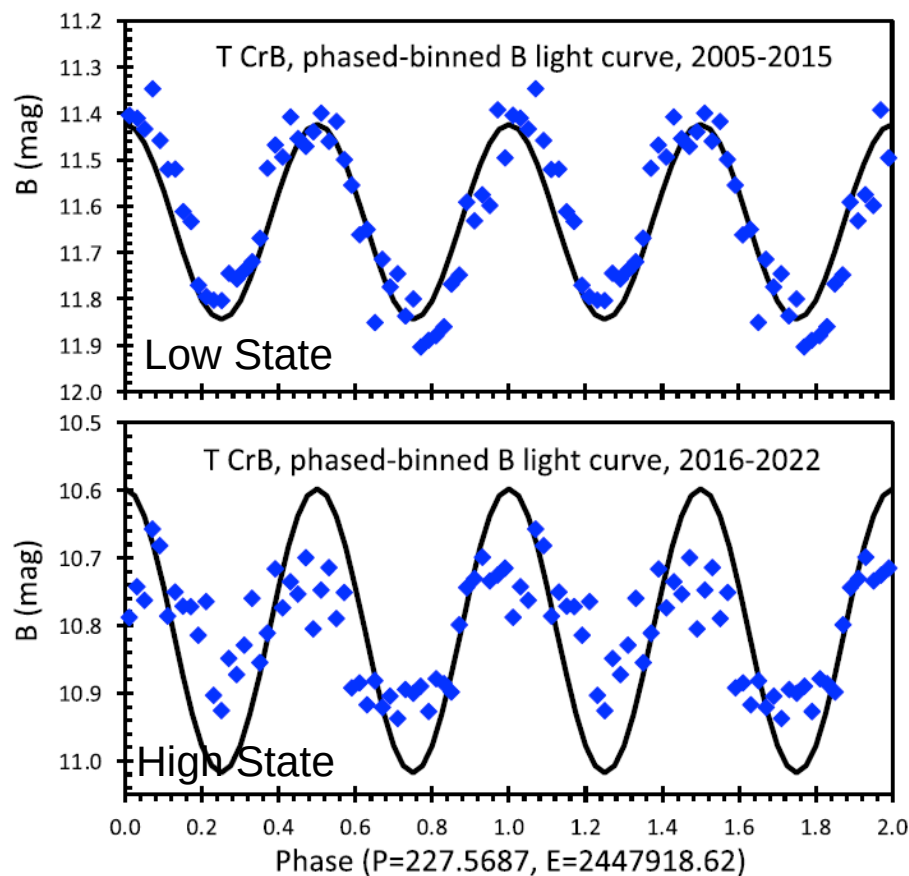
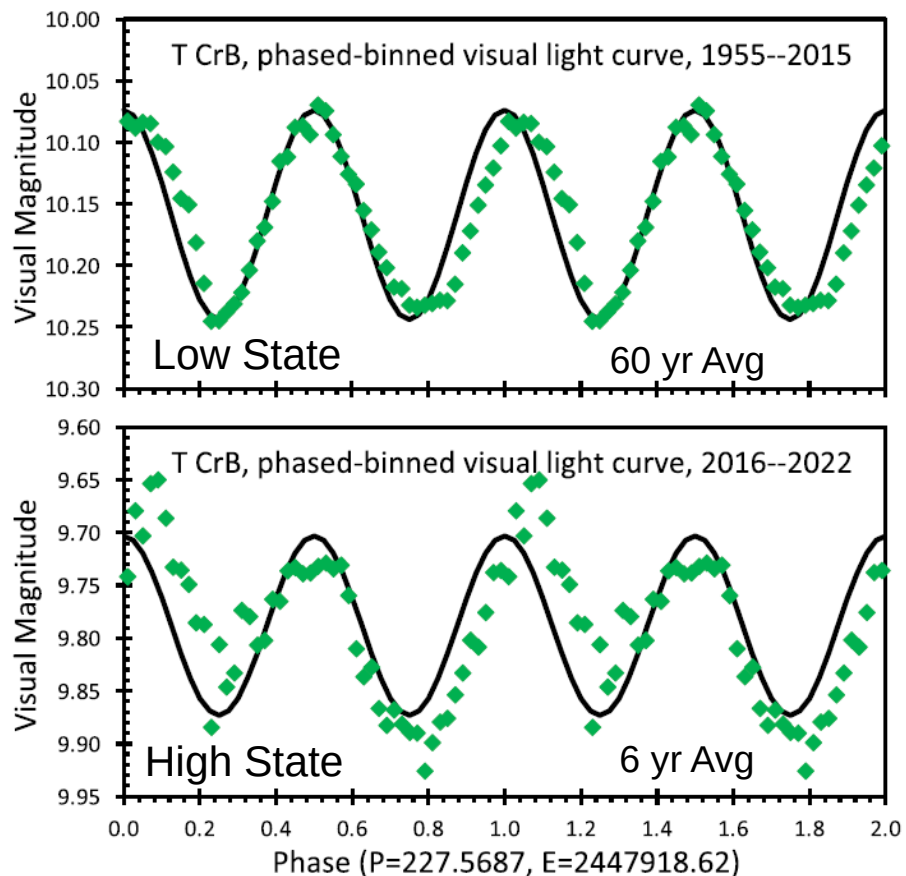
Pre-eruption Dip in B & V Apparent Magnitudes



Recent pre-eruption dip observed in 2015

B – Blue filter
V – Visual (green) filter

Orbit Phase Aligned Variability



How to see T CrB, the 'brightest nova of the generation'

- **Typically, T CrB is not visible to the naked eye at a magnitude of 10. When it reaches peak brightness, T CrB will be comparable to Polaris, the North Star, which has a magnitude of 2. (The smaller the number, the brighter it is.) It should remain this bright for a few days to a week before vanishing again. Astronomy.com**
- **Bootes and Hercules are relatively easy to find constellations – Corona Borealis is mid-way between them.**
- **No other stars except for the “crown” are bright enough to notice so T should stand out in dark area around constellation.**

The End

Until Next Cycle

HARDY

RS Oph

Illustration

Some References

- **Astronomical League - Binocular Observation Guide**
<https://nightsky.jpl.nasa.gov/docs/TCrB.pdf>
- **“The B&V light curves for recurrent nova T CrB from 1842-2022...”**, Schaefer, Bradley E., Dept Phy. & Astron., LSU 2023 Mar 5.
- How to see T CrB, the 'brightest nova of the generation', Elisabeth Gamillo, Astronomy.com, March 18, 2024.